On the concept of random orientation in far-field electromagnetic scattering by nonspherical particles

MICHAEL I. MISCHENKO¹,* AND MAXIM A. YURKIN²,³

¹NASA Goddard Institute for Space Studies, 2880 Broadway, New York, New York 10025, USA
²Voevodsky Institute of Chemical Kinetics and Combustion, SB RAS, Institutskaya Str. 3, 630090 Novosibirsk, Russia
³Novosibirsk State University, Pirogova 2, 630090 Novosibirsk, Russia
*Corresponding author: michael.i.mishchenko@nasa.gov

Received 24 November 2016; accepted 16 December 2016; posted 22 December 2016 (Doc. ID 281501); published 25 January 2017

Although the model of randomly oriented nonspherical particles has been used in a great variety of applications of far-field electromagnetic scattering, it has never been defined in strict mathematical terms. In this Letter, we use the formalism of Euler rigid-body rotations to clarify the concept of statistically random particle orientations and derive its immediate corollaries in the form of the most general mathematical properties of the orientation-averaged extinction and scattering matrices. Our results serve to provide a rigorous mathematical foundation for numerous publications in which the notion of randomly oriented particles and its light-scattering implications have been considered intuitively obvious. © 2017 Optical Society of America

OCIS codes: (290.5850) Scattering, particles; (290.5825) Scattering theory; (030.5620) Radiative transfer; (290.5855) Scattering, polarization.

https://doi.org/10.1364/OL.42.000494

The concept of randomly oriented nonspherical particles has been widely used in studies of far-field electromagnetic scattering and radiative transfer (see, e.g., Refs. [1–20] and references therein). Yet, the notion of random orientation has always been assumed to be self-explanatory, and its consequences such as optical isotropy of the extinction matrix and the dependence of the scattering matrix on only the scattering angle have been taken for granted as physically obvious and needing no proof. Furthermore, a specific integral over the three Eulerian angles has been assumed—again without explicit justification—to represent the desired averaging over random particle orientations [4,7–12,14–19,21,22].

One should keep in mind, however, that literature on electromagnetic scattering by particles contains many examples of "physical obviousness" turning out to be a shaky and misleading argument. Therefore, the great practical importance of the model of electromagnetic scattering by randomly oriented particles makes it essential to replace the presumed physical obviousness of the main consequences of this model with explicit mathematical proofs. We do that in the rest of this Letter.

Although the notion of a randomly oriented particle may appear to be intuitively evident, it needs to be formulated and parameterized mathematically before it can be used to derive useful corollaries. First, it is imperative to recognize that the randomness of particle orientations can be achieved only in the statistical sense, i.e., by assuming that (1) an optical observable is measured over a sufficiently long period of time and (2) the scattering process is ergodic [14,19]. These assumptions allow one to replace time averaging with ensemble averaging, including averaging over particle orientations. In this Letter, we apply the concept of random orientations to the far-field extinction and scattering matrices. As such, it is relevant to scattering by a single random particle or a small and sparse random multiparticle group as well as to the radiative transfer theory. For simplicity, we will consider only far-field scattering by a single random particle or a random two-particle group, but generalization to other relevant situations is straightforward [14,19].

Second, it is quite natural to parameterize the orientation of a particle with respect to a fixed right-handed reference frame Oxyz by affixing a right-handed reference frame Ox'y'z' to the particle and specifying the orientation of Ox'y'z' relative to Oxyz. Note that the two reference frames are assumed to have the same origin representing, e.g., the center of mass of the particle (Fig. 1).

It has been known since the classical work by Leonhard Euler that the orientation of Ox'y'z' with respect to Oxyz can be uniquely parameterized (except when β = 0 or π) by the set of three rotation angles $\mathbf{g} = \{\alpha, \beta, \gamma\}$, as shown in Fig. 2. It is therefore natural to parameterize the orientation of the particle by specifying the three Euler angles that transform the laboratory reference frame Oxyz into the particle reference frame Ox'y'z'. A statistical distribution of particle orientations can then be parameterized by the normalized probability density function $P(g)$ such that

$$\int_0^{2\pi} d\alpha \int_0^{2\pi} d\beta \int_0^{2\pi} d\gamma P(g) = 1. \quad (1)$$

The average of a function $f(g)$ over particle orientations is then calculated according to...
Two particle reference frames having a common origin.

The next step is to determine the form of the probability density function \( P_1(g) \) corresponding to the random (or uniform) orientation distribution. Let us affix to a nonspherical particle two different particle reference frames, \( O'x'y'z' \) and \( O''x''y''z'' \), having the same origin (Fig. 3). We will determine \( P_1(g) \) by requiring that irrespective of the particle morphology and of the choice of the two particle reference frames, averaging any function of particle orientation over the uniform orientation distribution of either \( O'x'y'z' \) or \( O''x''y''z'' \) yields exactly the same result. This natural definition is an essential step from a physical notion of random orientation to rigorous mathematical implications.

Let \( \tilde{g} = \{\tilde{a}, \tilde{b}, \tilde{c}\} \) be the rotation transforming \( O''x''y''z'' \) into \( O'x'y'z' \) and \( g \) be the rotation resulting from the rotation \( \tilde{g} \) followed by the rotation \( g = \{a, b, c\} \). Then, the above definition of randomness of particle orientations implies the identity

\[
\int_0^{2\pi} d\alpha \int_0^\pi d\beta \int_0^{2\pi} d\gamma f(\tilde{g}) P_1(g) = \int_0^{2\pi} d\alpha \int_0^\pi d\beta \int_0^{2\pi} d\gamma f(g) P_1(g),
\]

valid for any fixed \( \tilde{g} \). In the context of the group theory, \( P_1(g) \) corresponds to a normalized right Haar measure on the 3D rotation group. This measure is invariant to a right translation by an arbitrary group element. The left Haar measure is defined by replacing \( gg \) with \( \tilde{g}g \) on the right-hand side of Eq. (3).

The normalized right Haar measure is unique for any parametrization of the group [24,25]. In particular, a fundamental consequence of Eq. (3) derived in Section 1.3 of Chap. 1 in Ref. [26] and Section 1.6 of Chap. 1 in Ref. [27] (see also Ref. [28]) is that

\[
P_1(g) = \frac{\sin \beta}{8\pi^2}.
\]

Moreover, the left and right Haar measures coincide for the 3D rotation group [27,28] (which implies that Eq. (3) remains valid with the same \( P_1(g) \) upon replacing \( g \tilde{g} \) with \( \tilde{g}g \)), so we do not distinguish between them further on. The derivation of Eq. (4) was originally performed using the \( zxy \) notation of the Euler angles (named by the axes of consecutive rotations). However, it is also valid for any other (proper) notation in which the first and the last axes coincide, including the \( zyx \) notation of Fig. 2. By contrast, the Haar measure is more complicated for Tait–Bryan angles wherein all three rotation axes are different [28]. This justifies the conventional choice of the Euler angles for the calculation of orientation-averaged optical properties.

Thus, we can conclude that averaging over random particle orientations is represented by

\[
\langle f \rangle = \frac{1}{16\pi^2} \int_0^{2\pi} d\alpha \int_0^\pi d\beta \sin \beta \int_0^{2\pi} d\gamma f(\alpha, \beta, \gamma).
\]
the laboratory reference frame Oxz into Ox′y′z′ and Ox1y1z1, respectively.

The implication of these definitions is as follows:

\[ K(\hat{n}_1;g) = K(\hat{n};e) \Rightarrow K(\hat{n}_1;gg) = K(\hat{n};g), \]  

where the second argument denotes the particle orientation and \( e \) is the identity element of the rotation group (i.e., no rotation). In other words, simultaneous rotation of the incident beam and the particle does not change the extinction matrix. Then, Eqs. (3) and (6) imply

\[ \langle K(\hat{n}_1;g) \rangle \equiv \langle K(\hat{n};g) \rangle = K, \]  

where the averaging is defined by Eq. (5). In other words, the orientation-averaged extinction matrix is independent of the direction of propagation of the incident plane wave and of the choice of the reference plane used to define the Stokes parameters.

The derivation of the general structure of \( K \) is based on the consideration of pairs of reciprocal scattering configurations as defined in Ref. [2], i.e., those with the incident and scattering directions interchanged and inverted. The standard assumption has always been that for each particle orientation there is another one called reciprocal. However, this assumption in Section 5.22 of Ref. [2] was not accompanied by an explicit mathematical definition of statistically random orientations, which left it unclear whether the numbers of original and reciprocal orientations are always equal. Therefore, we instead exploit Eq. (3) and the fact that the reciprocal configuration can be obtained from the original one by a single rotation \( \tilde{g} \). The particular form of this rotation is known for arbitrary \( \hat{n} \) and \( \hat{n}' \) [2] but is unimportant for the following. Equation (3) then yields

\[ \langle K(\hat{n};gg) \rangle = \langle K(\hat{n};g) \rangle = \frac{1}{2} \langle K(\hat{n};g\tilde{g}) + K(\hat{n};g) \rangle. \]

This formula allows for the application of the reciprocity relation for the amplitude scattering matrix [29], which leads to the cancellation of the elements \( K_{13}, K_{31}, K_{24}, \) and \( K_{42} \). Furthermore, owing to Eq. (7), the extinction matrix must be invariant with respect to an arbitrary rotation of the reference plane around \( \hat{n} \), \( L(\eta)KL(-\eta) \equiv K \), where \( L(\eta) \) is the Stokes rotation matrix for the rotation angle \( \eta \) [12,14,19]. It is then straightforward to show that the resulting orientation-averaged extinction matrix has a highly symmetric structure given by

\[ K = \begin{bmatrix} K_{11} & 0 & 0 & K_{14} \\ 0 & K_{11} & K_{23} & 0 \\ 0 & -K_{23} & K_{11} & 0 \\ K_{14} & 0 & 0 & K_{11} \end{bmatrix} \]

and has only three independent elements.

In addition, let us assume that the nonspherical particle is accompanied by its mirror counterpart. The following derivation requires the second particle to be a mirror image of the first particle with respect to the fixed scattering plane Oxz (assumption 2 in Section 5.22 of Ref. [2]). We set the particle reference frames such that they form a mirror-symmetric configuration with respect to the Oxz plane in the default orientation. We then note that the reflection \( b \) with respect to the Oxz plane is a combination of the rotation \( g \) around the \( y \) axis by an angle \( \pi \) followed by the inversion \( i \) (\( b = ig \)), and \( i \) commutes with any rotation or reflection (e.g., Section 4.1 of Ref. [30]). Then, the orientation average of any property of the mirror particle is given by

\[ \langle f(\hat{n};g) \rangle = \langle f(\hat{n};g) \rangle = \langle f(\hat{n};g\tilde{g}) \rangle, \]

where \( f(\hat{n};g) = f(\hat{n};g) \), and we have used Eq. (3) as well as its counterpart with \( g \) and \( g \) interchanged. In other words, Eq. (10) proves that reflections and rotations effectively commute inside the orientation averaging. In application to \( K \) this implies that

\[ \langle K(\hat{n}_1;g) \rangle + \langle K(\hat{n};g) \rangle = \langle K(\hat{n}_1;bg) + K(\hat{n};g) \rangle, \]

i.e., the averaging is performed on pairs for which the amplitude scattering matrices are intimately related [2], thereby causing the cancellation of the remaining off-diagonal elements of \( K \). The latter is then diagonal and has only one independent element:

\[ K = C_{\text{ext}} \text{diag}[1, 1, 1, 1], \]

where \( C_{\text{ext}} \) is the extinction cross section. It is easily seen that Eq. (12) is also valid for a single randomly oriented particle with a plane of symmetry.

Finally, let us consider the most general properties of the orientation-averaged \( 4 \times 4 \) Stokes scattering matrix. We remind the reader that the scattering matrix \( F(\hat{n}',\hat{n}) \) for the incidence direction \( \hat{n} \) and the scattering direction \( \hat{n}' \) relates the Stokes parameters of the incident plane wave and the scattered spherical wave in the far zone of the particle. In this case, both sets of the Stokes parameters are defined with respect to the scattering plane (i.e., the plane through \( \hat{n} \) and \( \hat{n}' \)).

As before, let \( \hat{n} \) and \( \hat{n}_1 \) be the unit vectors specifying two different incidence directions, and let the corresponding scattering planes be given by the Oxz and Ox1y1z1 planes of the reference frames Oxz and Ox1y1z1, respectively (Fig. 4). The corresponding scattering directions, \( \hat{n}' \) and \( \hat{n}_1' \), are obtained by rotating the vectors \( \hat{n} \) and \( \hat{n}_1 \) clockwise around the \( y \) and \( y_1 \) axis, respectively, through the same angle \( \Theta \) called the scattering angle.

Completely analogous to the case of the extinction matrix, simple geometrical considerations suggest that

\[ F(\hat{n}_1',\hat{n}_1;gg) = F(\hat{n}_1',\hat{n};g), \]

which, together with Eq. (3), implies that

\[ \langle F(\hat{n}_1',\hat{n}_1;g) \rangle = \langle F(\hat{n}_1',\hat{n};g) \rangle = F(\Theta), \]

Fig. 4. Laboratory, Oxz, and auxiliary, Ox1y1z1, reference frames.
where the averaging is defined by Eq. (5). Thus, the orientation-averaged scattering matrix is independent of the incidence direction and of the orientation of the scattering plane and is a function of only the scattering angle.

One can then use the reciprocity relation for the amplitude scattering matrix [29] and the line of thought in Section 5.22 of Ref. [2] to show that the orientation-averaged scattering matrix has 10 independent elements and is given by

\[
F(\Theta) = \begin{bmatrix}
F_{11}(\Theta) & F_{12}(\Theta) & F_{13}(\Theta) & F_{14}(\Theta) \\
F_{12}(\Theta) & F_{22}(\Theta) & F_{23}(\Theta) & F_{24}(\Theta) \\
-F_{13}(\Theta) & -F_{23}(\Theta) & F_{33}(\Theta) & F_{34}(\Theta) \\
F_{14}(\Theta) & F_{24}(\Theta) & -F_{34}(\Theta) & F_{44}(\Theta)
\end{bmatrix}.
\] (15)

The only essential missing step in Ref. [2] is the explicit use of Eq. (8) with \(K(\hat{n}, g)\) being replaced by \(F(\hat{n}, \hat{n}; g)\).

Also, assuming that the randomly oriented nonspherical particle is accompanied by its randomly oriented mirror counterpart and using Eq. (10) or the corresponding analog of Eq. (11) (with reflection defined with respect to the scattering plane) ultimately yields the following conventional block-diagonal scattering matrix (cf. Refs. [1,2]):

\[
F(\Theta) = \begin{bmatrix}
F_{11}(\Theta) & 0 & 0 & 0 \\
0 & F_{22}(\Theta) & 0 & 0 \\
0 & 0 & F_{33}(\Theta) & 0 \\
0 & 0 & 0 & F_{44}(\Theta)
\end{bmatrix}.
\] (16)

It can easily be shown that this formula also applies to a single randomly oriented particle with a plane of symmetry.

The previous discussion has been based on the assumption that all rotations of a particle occur in such a way that a point inside the particle (e.g., its center of mass) remains fixed at the origin of the laboratory reference frame. However, the extinction and scattering matrices are invariant with respect to parallel translations of the particle as long as these translations are sufficiently small (see Section 3.1 of Ref. [12] or Section 13.9 of Ref. [19]). Therefore, all properties of the orientation-averaged extinction and scattering matrices remain the same if the particle wobbles arbitrarily during the measurement.

In summary, we began this Letter by giving the explicit mathematical definition of a randomly oriented (in the statistical sense) nonspherical particle. We then derived the main properties of the far-field extinction and scattering matrices as direct mathematical corollaries of Eq. (3) rather than assuming them to be intuitively obvious. As a result, we have provided the rigorous (and long overdue) justification of the main premises implicit in Refs. [1–20] and numerous related publications.

**Funding.** National Aeronautics and Space Administration (NASA) (Remote Sensing Theory Program and ACE Project); Russian Science Foundation (RSF) (14-15-00155).

**Acknowledgment.** We appreciate instructive discussions with Joop Hovenier, Andrey Mishchenko, Katharina Sander, and Larry Travis as well as two constructive anonymous reviews.

**REFERENCES**